

Compilation of recent nuclear ground state charge radius measurements and tests for models

Tao Li ^{a,b,*}, Yani Luo ^a, Ning Wang ^{a,b,**}

^a Department of Physics, Guangxi Normal University, Guilin, 541004, PR China

^b Guangxi Key Laboratory of Nuclear Physics and Technology, Guilin, 541004, PR China



ARTICLE INFO

Article history:

Received 21 January 2021

Received in revised form 17 April 2021

Accepted 19 April 2021

Available online 7 May 2021

Keywords:

Nuclear charge radius

Shell effect

Mirror nuclei

Isospin asymmetry

ABSTRACT

The root-mean-square (rms) charge radii of 236 nuclides measured by laser spectroscopy experiment are compiled, and the uncertainties are calculated. From the rms charge radii of Mg isotope chain, the new magic number $N = 14$ can be observed, and the traditional magic number $N = 20$ disappears in the K isotope chain. A good linear relationship between the difference of the mirror nuclear charge radii and the isospin asymmetry can be clearly observed with a Pearson's linear correlation coefficient of 0.96. The accuracies and predictive powers of the WS* and HFB25 models are tested with these new data. The rms deviation with the WS* model is only 0.0176 fm for the 129 new data, which is much better than that of HFB25 model. The influence of deformation effect on nuclear charge radius prediction is studied systematically.

© 2021 Elsevier Inc. All rights reserved.

* Corresponding author at: Department of Physics, Guangxi Normal University, Guilin, 541004, PR China.

** Corresponding author.

E-mail addresses: litao@gxnu.edu.cn (T. Li), [\(N. Wang\)](mailto:wangning@gxnu.edu.cn).

Contents

| | |
|--|---|
| 1. Introduction..... | 2 |
| 2. New data on nuclear charge radii..... | 2 |
| 3. Tests of the theoretical methods for nuclear charge radii..... | 3 |
| 4. Summary | 4 |
| Declaration of competing interest..... | 5 |
| Acknowledgments | 5 |
| References | 5 |
| Explanation of Tables..... | 7 |
| Table 1. The experimental data of rms charge radii for 236 nuclei..... | 7 |

Table 1. The experimental data of rms charge radii for 236 nuclei.

1. Introduction

As one of the static properties in atomic nuclei, nuclear charge radius is a key observable that can directly reflect the important characteristics on the nuclear structure. For instance, nuclear charge radius could give signals for the occurrence of new magic number or the disappearance of traditional magic number considering the influence of shell effect on charge radii [1–3]. In addition, it is useful for the study of neutron skin, neutron halo and isomer [4–6]. So far, the root-mean-square (rms) charge radii of more than 1000 nuclei have been measured by two types of experiments in general: (i) the charge radii of stable nuclei were measured by charged particle scattering experiment, (ii) the charge radii of radionuclides were measured by charge radii changes $\delta\langle r^2 \rangle$ extracted from laser spectroscopy and K_α X-ray isotope shifts [7,8]. In recent years, with the development of experiment technology, more and more rms charge radii of unstable nuclei have been measured for the first time by charge radii changes $\delta\langle r^2 \rangle$ [9–35] from laser spectroscopy experiments. It is therefore interesting to systematically study the new experimental data.

In the beginning, the nuclear charge radius R_0 is usually described by the $A^{1/3}$ law: $R_0 = r_0 A^{1/3}$, where A is the mass number. With more experimental data being obtained, it was found that the isospin and shell effects also play very important roles for the charge radius [2,36–41]. The rms nuclear charge radius can be self-consistently calculated by using microscopic nuclear mass models, such as the Skyrme–Hartree–Fock–Bogoliubov (HFB) model [42,43] and the relativistic mean field (RMF) model [44,45]. In addition, the nuclear charge radius can also be predicted by using some local relations [46–48] such as the Garvey–Kelson relations (GKRs). It is necessary to test the predictive power of these different models for the description of nuclear charge radius based on new measured data. In addition, nuclear rms charge radius is closely related to deformation parameters. It is therefore interesting to study the influence of deformation parameters on the calculation of nuclear rms charge radius.

The structure of this paper is as follows: In Section 2, 236 experimental data for nuclear rms charge radii from laser spectroscopy experiments are compiled and the corresponding uncertainties are analyzed. The influence of shell effect on the charge radius will be also investigated for Mg and K isotope chains in this section. In Section 3, the accuracy and predictive power of the two models will be tested based on these new experimental data, and the influence of different deformation parameters on the calculation of nuclear rms charge radii will be analyzed. Finally, a summary will be given in Section 4.

2. New data on nuclear charge radii

The mean square charge radius difference $\delta\langle r^2 \rangle^{A'A} = \langle r^2 \rangle^A - \langle r^2 \rangle^{A'}$ between isotopes can be obtained from the isotope shift based on laser spectroscopy experiment. Therefore, mean square

charge radius can be written as $\langle r^2 \rangle^A = \langle r^2 \rangle^{A'} + \delta\langle r^2 \rangle^{A'A}$. The rms charge radius can be calculated from the following formula

$$r_c(A) = \sqrt{\langle r^2 \rangle^A} = \sqrt{\langle r^2 \rangle^{A'} + \delta\langle r^2 \rangle^{A'A}} \quad (1)$$

where A' represents the mass number of stable reference isotope. According to the calculation method of uncertainty, the uncertainty of nuclear rms charge radius from experiment can be expressed as

$$\Delta r_c(A) = r_c(A) \cdot \frac{\sqrt{4r_c^2(A') \cdot (\Delta r_c(A'))^2 + (\Delta\delta\langle r^2 \rangle^{A'A})^2}}{2[r_c^2(A') + \delta\langle r^2 \rangle^{A'A}]} \quad (2)$$

where $r_c(A')$ and $\Delta r_c(A')$ are the rms charge radius and uncertainty of A' isotope, $\Delta\delta\langle r^2 \rangle^{A'A}$ is the uncertainty of the mean square charge radius difference $\delta\langle r^2 \rangle^{A'A}$ and extracted from experimental statistical and systematic errors. Usually, the literature will present the corresponding statistical and systematic errors as well as the mean square charge radius difference of isotopes. Using Eqs. (1) and (2), we obtain 236 experimental data of nuclear charge radius measured by laser spectroscopy experiment and calculate the corresponding uncertainties, which are listed in Table 1, where $r_c(A')$ is taken from Ref. [1] and has been set in bold in Table 1. Compared with the data in Ref. [1], there are 130 new experimental data. It is worth noting that the rms charge radii of 106 nuclei in Table 1 are slightly different from the published results in Ref. [1], which may be caused by the fact that the original experimental data were selected from different experimental groups or different experimental measurement methods were adopted. We checked the uncertainties calculated by Eq. (2) and found that our results for Be isotope are in agreement with the results in Ref. [9]. Fig. 1 shows the uncertainties of rms charge radius for $Z = 29$ and 48 isotope chains calculated by Eq. (2). One can see from Fig. 1, the uncertainties obviously increase with the distance from isotope A' (dotted line position). The uncertainties of rms charge radius of other isotopic chains measured by laser spectroscopy have similar characteristics.

The nuclear shell effect is one of the important microscopic quantum effects in nuclear physics. We will focus on the effect of neutron shell on charge radius in isotope chains. Similar to neutron shell gap of nuclear mass, we define the neutron gap of charge radius as follows:

$$\Delta_n^{r_c}(N, Z) = r_c(N - 2, Z) + r_c(N + 2, Z) - 2r_c(N, Z) \quad (3)$$

where $r_c(N, Z)$ is the measured nuclear rms charge radii. In Fig. 2, we show the trend of neutron gaps of charge radii of Mg and K isotope chains. For K isotope chain, one can clearly see the peak at the magic number $N = 28$. However, the enhancement of $\Delta_n^{r_c}$ cannot be observed at the magic number $N = 20$. A similar behavior can also be observed in the Ar and Ca isotope chains. It is known that the doubly magic nuclei are spherical in shape, the abrupt change trend of the charge radius at the magic number could be due to the shell effect in nucleus. For Mg isotope chain, the peak at $N = 14$ can be clearly observed in the trend of

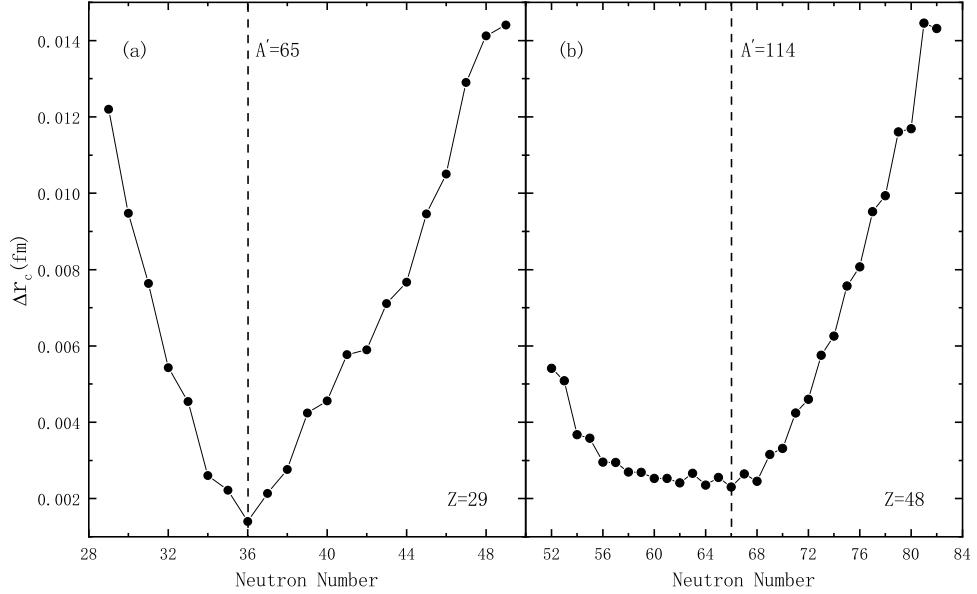


Fig. 1. Uncertainties of rms charge radii for $Z = 29$ and $Z = 48$ isotope chains are calculated by Eq. (2).

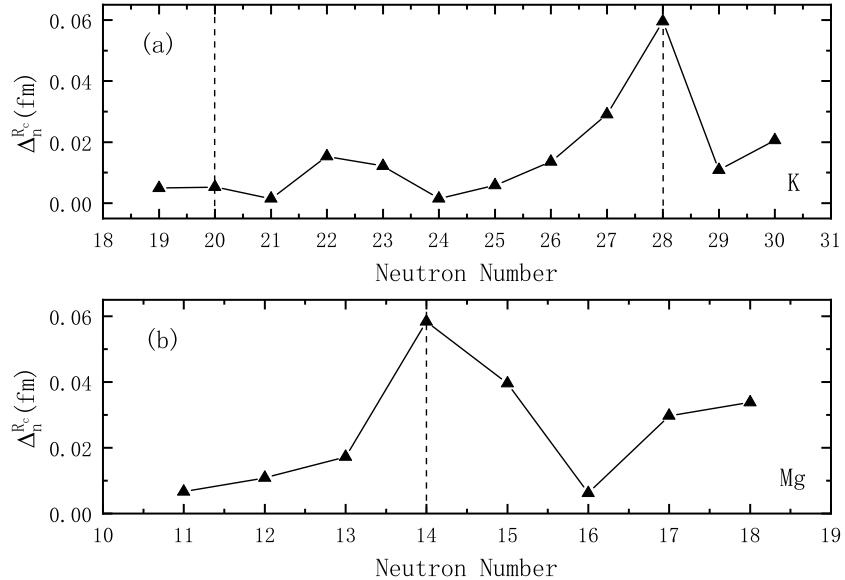


Fig. 2. Neutron gaps of charge radii of K and Mg isotope chains as a function of neutron number.

neutron gaps of charge radii, which implies that $N = 14$ could be a new magic number. In addition, the evidence that $N = 14$ is a magic number can also be seen from the trend of the two-neutron separation energy with the neutron number. Recently, S. Bagchi et al. also found the evidence for magic number $N = 14$ from the proton radii of N isotope chain [49].

In the absence of coulomb interactions between protons, a perfectly charge-symmetric and charge-independent nuclear force would result in the binding energies of mirror nuclei (i.e. nuclei with the same atomic number A but with the proton number Z and neutron number N interchanged) being identical. The neutron skin thickness of a neutron-rich nucleus should be very close to the proton skin thickness of its mirror nucleus, if we remove the coulomb interactions. Considering the difficulties in the measurement of neutron skin thickness, it is interesting to study the difference of charge radii between mirror nuclei. We study the relationship between the difference in mirror nuclei charge radii and the isospin asymmetry $I = (N - Z)/A$. A good

linear relationship between them can be observed from Fig. 3, and Pearson's linear correlation coefficient $P = 0.96$. It indicates that the difference of mirror nuclear charge radii could be useful for the study of symmetry energy at saturated densities [2].

3. Tests of the theoretical methods for nuclear charge radii

In Ref. [2], considering the isospin and shell effects in the nucleus, a four-parameter nuclear charge radii formula was proposed by combining the shell corrections and deformations of nuclei obtained from the Weizsäcker-Skyrme (WS*) mass model [50]. The expression is shown below

$$r_c = \sqrt{\frac{3}{5}} [1.226A^{1/3} + 2.86A^{-2/3} - 1.09I(1-I) + 0.99\Delta E/A] \times [1 + \frac{5}{8\pi}(\beta_2^2 + \beta_4^2)] \quad (4)$$

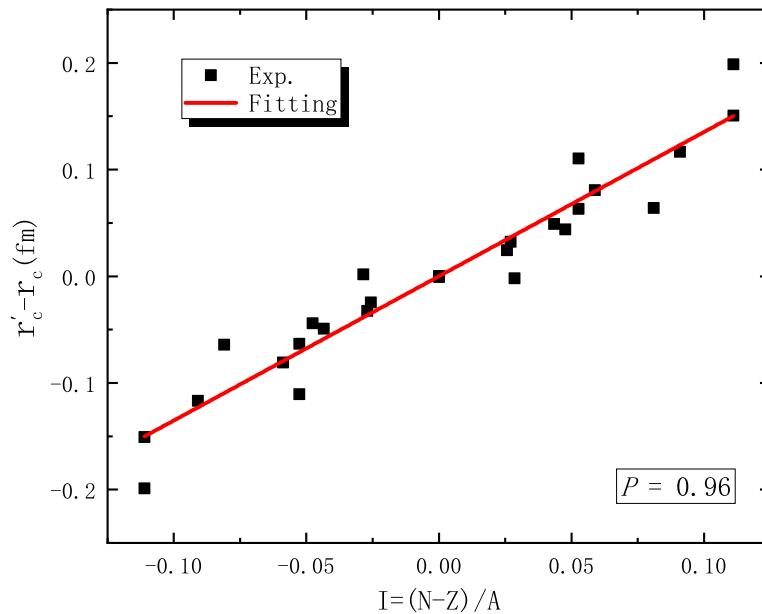


Fig. 3. (Color online) Linear relationship between the difference of mirror nuclei charge radii and isospin asymmetry I , where r_c and r'_c represent the rms charge radii of nuclei and its corresponding mirror nuclei, respectively.

where A is the mass number, I is the isospin asymmetry, ΔE denotes the shell corrections, β_2 and β_4 represent the quadrupole and hexadecapole deformation of the nucleus, respectively. The 885 measured rms charge radii of the nuclei from Ref. [1] were reproduced by this formula with an rms deviation $\sigma = \sqrt{\frac{1}{n} \sum_{i=1}^n (r_c^{exp} - r_c^{th})^2} = 0.022$ fm. In addition, Hartree-Fock-Bogoliubov (HFB) mass models [42,43] were successfully proposed for describing nuclear properties, such as nuclear mass and charge radius. The corresponding rms deviation with respect to the 884 measured charge radii is 0.025 fm from the HFB25 model calculations [42], which is comparable to the results of WS* model (Eq. (4)). It is interesting to test the predictive power of these two models based on the latest experimental data of nuclear charge radius.

Fig. 4(a) shows the deviation between the measured rms charge radii for 1014 nuclei and model predictions. The rms deviations are 0.0213 fm from WS* model and 0.0254 fm from the HFB25 model, respectively. In Table 1, we list the predictions of these two models for 129 new measured charge radii. Fig. 4(b) shows the difference between the experimental data and model predictions for the 129 new charge radii. One can see results of the WS* model are much better than those of HFB25 model.

In addition to the difference in describing the ground state mass of nucleus, it is also different for the HFB25 and WS* models in describing the ground state charge radius of nucleus. Literature [51] also shows that the deformation energies of the HFB25 model are systematically larger than those of the WS* model, especially for light and intermediate nuclei. Since it is difficult to directly measure the nuclear deformation parameters experimentally, it is necessary to study the influence of deformation parameters on the prediction of nuclear charge radius. We introduce a correction coefficient r_β of the deformation parameter in Eq. (4), and the expression is as follows.

$$r_c = \sqrt{\frac{3}{5}} [r_0 A^{1/3} + r_1 A^{-2/3} + r_s I(1-I) + r_d \Delta E/A] \left[1 + \frac{5r_\beta}{8\pi} (\beta_2^2 + \beta_4^2) \right] \quad (5)$$

The introduction of the coefficient r_β may further improve the predictions of the rms charge radius due to the inaccuracy of

the deformation parameters given by the theoretical model. Of course, it also checks the overall behavior of the deformation parameters given by the theoretical model, because if the deformation parameters are extremely accurate, the coefficient r_β should be strictly equal to 1.

The coefficients in Eq. (5) can be obtained by fitting the latest 1014 experimental data of atomic nuclei rms charge radii based on the shell correction energies and deformation parameters provided by different nuclear mass models, which are listed in Table A. The data in the first and second rows in Table A are obtained by fitting the latest experimental data when the deformation correction coefficient $r_\beta = 1$, which is actually equivalent to the new parameters in Eq. (4). First of all, it can be seen from the third column of Table A that although a deformation correction coefficient is added to the formula, the rms deviation is only reduced by 2.3% with the shell correction energies and deformation parameters provided by the WS* mass model. However, the rms deviation obtained by fitting the experimental data with deformation parameters provided by the HFB25 mass model is reduced by 11.2%. It indicates that deformation parameters significantly influence the calculation of nuclear charge radius. It is worth mentioning that the optimally obtained coefficients in Eq. (5) and rms deviation based on the deformation parameters and shell corrections from the latest WS4 mass model [52] are very close to those from the WS* mass model. From the value of r_β in Table A, it seems that the deformation parameters given by the WS* mass model are underpredicted, while the deformation parameters given by the HFB25 mass model are overpredicted. For the prediction of charge radii of deformed nuclei, the formula in Eq. (5) could give better results.

4. Summary

In this work, we compiled the rms charge radii of 236 nuclides measured by laser experiment, and systematically calculated the uncertainties. The new magic number $N = 14$ was observed in the Mg isotope chain based on the neutron gaps of charge radii, and the traditional magic number $N = 20$ disappeared in the K isotope chain. A good linear relationship between the difference of the mirror nuclear charge radii and the isospin asymmetry

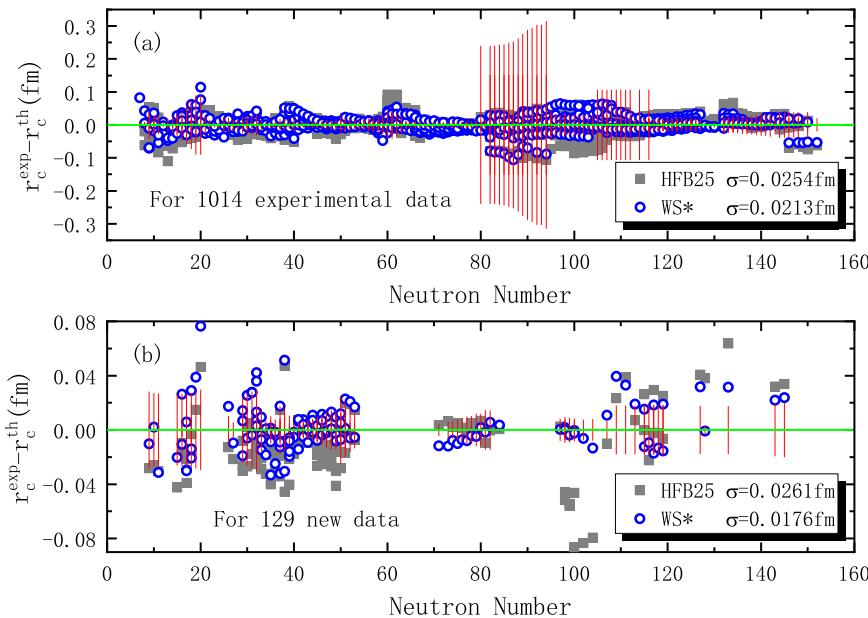


Fig. 4. Difference between the experimental data and model calculations for the rms nuclear charge radius. The circles and squares denote the results of the WS* and HFB25 models, respectively. The red error bars denote the uncertainty of the experimental data. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

Table A

Parameters in Eq. (5) obtained by fitting latest experimental data, the shell correction energies from the WS* mass model and the deformation parameters from the WS* and HFB25 mass models. The first column indicates which model the deformation parameters are taken from. The second column represents the number of experimental data required for fitting. The third column represents the rms deviation.

| Model | N_{nucl} | σ (fm) | r_0 | r_1 | r_s | r_d | r_β |
|-------|------------|---------------|-------|-------|--------|-------|-----------|
| WS* | 1014 | 0.0214 | 1.225 | 2.884 | -1.066 | 0.948 | 1 |
| HFB25 | 1013 | 0.0286 | 1.217 | 3.023 | -0.972 | 0.647 | 1 |
| WS* | 1014 | 0.0209 | 1.224 | 2.916 | -1.086 | 0.920 | 1.213 |
| HFB25 | 1013 | 0.0254 | 1.222 | 2.903 | -0.963 | 0.830 | 0.600 |

was clearly observed with a Pearson's linear correction coefficient of 0.96. We tested the accuracies and predictive powers of the WS* and HFB25 models based on the new experimental data of charge radius. The rms deviation between predictions from the WS* model and 1014 known charge radii is 0.0213 fm, and the rms deviation between predictions from the HFB25 model and 1013 known charge radii is 0.0254 fm. The rms deviation between the predicted results of the WS* model and the 129 new experimental value is only 0.0176 fm, smaller by 20% than the results with respect to the 885 experimental data. The corresponding result for the HFB25 model is 0.0261 fm, which is larger by 4.4% than the previous results of the 884 experimental data. In addition, we find that deformation parameters have an obvious influence on the prediction of nuclear charge radius. For a better consideration of the influence of deformation parameters on the prediction of nuclear charge radius, the charge radii formula is revised. The deformation parameters from the WS* and HFB25 mass models are roughly evaluated by using Eq. (5) and newest 1014 experimental data of nuclear rms charge radii. We find that the deformation parameters given by the WS* mass model are slightly underpredicted, while those given by HFB25 mass model are overpredicted.

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

Acknowledgments

This work was supported by the National Natural Science Foundation of China (Grant No. U1867212, 12047567), the Natural Science Foundation of Guangxi (Grant No. 2017GXNSFGA198001) and the Middle-aged and Young Teachers' Basic Ability Promotion Project of Guangxi (CN) (Grant No. 2019KY0061).

References

- [1] I. Angeli, K.P. Marinova, At. Data Nucl. Data Tables 99 (2013) 69.
- [2] N. Wang, T. Li, Phys. Rev. C 88 (2013) 011301, (R).
- [3] I. Angeli, K.P. Marinova, J. Phys. G: Nucl. Part. Phys. 42 (2015) 055108, 20pp.
- [4] W. Geithner, T. Neff, G. Audi, et al., Phys. Rev. Lett. 101 (2008) 252502.
- [5] W. Nörtershäuser, D. Tiedemann, M. Žáková, et al., Phys. Rev. Lett. 102 (2009) 062503.
- [6] D.T. Yordanov, D.L. Balabanski, M.L. Bissell, et al., Phys. Rev. Lett. 116 (2016) 032501.
- [7] K. Marinova, J. Phys. Chem. Ref. Data 44 (3) (2015) 031214.
- [8] I. Angeli, K. Marinova, J. Phys. Conf. Ser. 724 (2016) 012032.
- [9] A. Krieger, K. Blaum, M.L. Bissell, et al., Phys. Rev. Lett. 108 (2012) 142501.
- [10] D.T. Yordanov, M.L. Bissell, K. Blaum, et al., Phys. Rev. Lett. 108 (2012) 042504.
- [11] D.M. Rossi, K. Minamisono, H.B. Asberry, et al., Phys. Rev. C 92 (2015) 014305.
- [12] A.J. Miller, K. Minamisono, A. Klose, et al., Nat. Phys. 15 (2019) 432.
- [13] R.F. Garcia Ruiz, M.L. Bissell, K. Blaum, et al., Nat. Phys. 12 (2016) 594.
- [14] H. Heylen, C. Babcock, R. Beerwerth, et al., Phys. Rev. C 94 (2016) 054321.
- [15] K. Minamisono, D.M. Rossi, R. Beerwerth, et al., Phys. Rev. Lett. 117 (2016) 252501.
- [16] S. Kaufmann, J. Simonis, S. Baccia, et al., Phys. Rev. Lett. 124 (2020) 132502.
- [17] M.L. Bissell, T. Carette, K.T. Flanagan, et al., Phys. Rev. C 93 (2016) 064318.
- [18] R.P. de Groot, J. Billowes, C.L. Binnersley, et al., Nat. Phys. 16 (2020) 620.
- [19] L. Xie, X.F. Yang, C. Wraith, et al., Phys. Lett. B 797 (2019) 134805.
- [20] T.J. Procter, J. Billowes, M.L. Bissell, et al., Phys. Rev. C 86 (2012) 034329.
- [21] G.J. Farooq-Smith, A.R. Vernon, J. Billowes, et al., Phys. Rev. C 96 (2017) 044324.
- [22] E. Mané, A. Voss, J.A. Behr, et al., Phys. Rev. Lett. 107 (2011) 212502.
- [23] T.J. Procter, J.A. Behr, J. Billowes, et al., Eur. Phys. J. A 51 (2015) 23.
- [24] R. Ferrer, N. Bree, T.E. Cocolios, et al., Phys. Lett. B 728 (2014) 191.
- [25] M. Hammen, W. Nörtershäuser, D.L. Balabanski, et al., Phys. Rev. Lett. 121 (2018) 102501.
- [26] C. Gorges, L.V. Rodríguez, D.L. Balabanski, et al., Phys. Rev. Lett. 122 (2019) 192502.

- [27] K.T. Flanagan, J. Billowes, P. Campbell, et al., *J. Phys. G: Nucl. Part. Phys.* 39 (2012) 125101.
- [28] B.A. Marsh, T. Day Goodacre, S. Sels, et al., *Nat. Phys.* 14 (2018) 1163.
- [29] A.E. Barzakh, A.N. Andreyev, T.E. Cocolios, et al., *Phys. Rev. C* 95 (2017) 014324.
- [30] A.E. Barzakh, L. Kh. Batist, D.V. Fedorov, et al., *Phys. Rev. C* 88 (2013) 024315.
- [31] A.E. Barzakh, D.V. Fedorov, V.S. Ivanov, et al., *Phys. Rev. C* 97 (2018) 014322.
- [32] M.D. Seliverstov, T.E. Cocolios, W. Dexters, et al., *Phys. Lett. B* 719 (2013) 362.
- [33] D.A. Fink, T.E. Cocolios, A.N. Andreyev, et al., *Phys. Rev. X* 5 (2015) 011018.
- [34] K.M. Lynch, J. Billowes, M.L. Bissell, et al., *Phys. Rev. X* 4 (2014) 011055.
- [35] K.M. Lynch, S.G. Wilkins, J. Billowes, et al., *Phys. Rev. C* 97 (2018) 024309.
- [36] Bożena Nerlo-Pomorska, Z. Krzysztof Pomorski, *Phys. A* 344 (1993) 359.
- [37] Bożena Nerlo-Pomorska, Z. Krzysztof Pomorski, *Phys. A* 348 (1994) 169.
- [38] I. Angeli, N.S. Aph, *Heavy Ion Phys.* 13 (2001) 149.
- [39] I. Angeli, *At. Data Nucl. Data Tables* 87 (2004) 185.
- [40] R.H. Li, Y.M. Hu, M.C. Li, *Physica C* 33 (2009) 123, (Suppl. 1).
- [41] Z.Q. Sheng, G.W. Fan, J.F. Qian, et al., *Eur. Phys. J. A* 51 (2015) 40.
- [42] S. Goriely, N. Chamel, J.M. Pearson, *Phys. Rev. C* 88 (2013) 024308.
- [43] S. Goriely, N. Chamel, J.M. Pearson, *Phys. Rev. C* 88 (2013) 061302, (R).
- [44] G.A. Lalazissis, S. Raman, P. Ring, *At. Data Nucl. Data Tables* 71 (1999) 1.
- [45] P.W. Zhao, Z.P. Li, J.M. Yao, et al., *Phys. Rev. C* 82 (2010) 054319.
- [46] J. Piekarewicz, M. Centelles, X. Roca-Maza, et al., *Eur. Phys. J. A* 46 (2010) 379.
- [47] B.H. Sun, Y. Lu, J.P. Peng, et al., *Phys. Rev. C* 90 (2014) 054318.
- [48] M. Bao, Y. Lu, Y.M. Zhao, et al., *Phys. Rev. C* 94 (2016) 064315.
- [49] S. Bagchi, R. Kanungo, W. Horiuchi, et al., *Phys. Lett. B* 790 (2019) 251.
- [50] N. Wang, Z.Y. Liang, M. Liu, et al., *Phys. Rev. C* 82 (2010) 044304.
- [51] N. Wang, T. Li, *Acta Phys. B. Pro. Sup.* 12 (2019) 715.
- [52] N. Wang, M. Liu, X.Z. Wu, et al., *Phys. Lett. B* 734 (2014) 215.

Explanation of Tables

Table 1. The experimental data of rms charge radii for 236 nuclei

| | |
|-----------------------------------|---|
| Z | The atomic (proton) number. |
| el. | The element symbol. |
| N | The neutron number. |
| A | The mass number. |
| $\delta\langle r^2 \rangle$ | The mean square charge radius difference comes from the corresponding references, as shown below. |
| $\Delta\delta\langle r^2 \rangle$ | The error of the mean square charge radius difference $\Delta\delta\langle r^2 \rangle = \sqrt{\Delta_{\text{sta.}}^2 + \Delta_{\text{sys.}}^2}$, where $\Delta_{\text{sta.}}$ and $\Delta_{\text{sys.}}$ represent the statistical error and systematic error of $\delta\langle r^2 \rangle$ respectively, from the corresponding references. |
| r_c | The experimental data of nuclear rms charge radius, it is calculated by Eq. (1). |
| Δr_c | The uncertainty of nuclear rms charge radius, it is calculated by Eq. (2). |
| In Ref. [1] | Specifies if the nuclear charge radius measurement was included in Ref. [1] |
| (1) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{7.9-12}\text{Be}$ from Ref. [9]. |
| (2) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{21-32}\text{Mg}$ from Ref. [10]. |
| (3) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{36-51}\text{K}$ from Ref. [11]. |
| (4) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{36-39}\text{Ca}$ and $^{43-52}\text{Ca}$ from Refs. [12] and [13], respectively. |
| (5) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{50-64}\text{Mn}$ from Ref. [14]. |
| (6) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{52,53}\text{Fe}$ from Ref. [15]. |
| (7) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{58,61,62,64,68}\text{Ni}$ from Ref. [16]. |
| (8) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{58-75}\text{Cu}$ and $^{76-78}\text{Cu}$ from Refs. [17] and [18], respectively. |
| (9) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{62-80}\text{Zn}$ from Ref. [19]. |
| (10) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{63-82}\text{Ga}$ from Refs. [20,21]. |
| (11) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{74-76}\text{Rb}$ from Refs. [22,23]. |
| (12) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{97-101,107}\text{Ag}$ from Ref. [24]. |
| (13) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{100-130}\text{Cd}$ from Ref. [25]. |
| (14) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of even- A $^{108-134}\text{Sn}$ from Ref. [26]. |
| (15) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{175,177}\text{Yb}$ from Ref. [27]. |
| (16) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{177-185}\text{Hg}$ from Ref. [28]. |
| (17) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{179-181,183}\text{Tl}$ and $^{185,190-205,207,208}\text{Tl}$ from Refs. [29] and [30], respectively. |
| (18) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{211,213}\text{Bi}$ from Ref. [31]. |
| (19) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of odd- A $^{193-203,209-211}\text{Po}$ and $^{217,218}\text{Po}$ from Refs. [32] and [33], respectively. |
| (20) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{202-206}\text{Fr}$ from Ref. [34]. |
| (21) | The mean square charge radius differences $\delta\langle r^2 \rangle$ of $^{222-233}\text{Ra}$ from Ref. [35]. |

Table 1

The experimental data of rms charge radii for 236 nuclei.

| Z | el. | N | A | $\delta(r^2)$ (fm 2) | $\Delta\delta(r^2)$ (fm 2) | r_c (fm) | Δr_c (fm) | In Ref. [1] |
|-----|-----|-----------|-----------|--------------------------|--------------------------------|---------------|-------------------|-------------|
| 4 | Be | 3 | 7 | 0.66 | 0.06 | 2.6468 | 0.0161 | Yes |
| | | 5 | 9 | 0 | – | 2.5190 | 0.0120 | Yes |
| | | 6 | 10 | −0.77 | 0.05 | 2.3612 | 0.0166 | Yes |
| | | 7 | 11 | −0.26 | 0.04 | 2.4669 | 0.0147 | Yes |
| | | 8 | 12 | −0.08 | 0.05 | 2.5031 | 0.0157 | No |
| 12 | Mg | 9 | 21 | 0.176 | 0.068 | 3.0629 | 0.0114 | No |
| | | 10 | 22 | 0.214 | 0.052 | 3.0691 | 0.0089 | No |
| | | 11 | 23 | 0.053 | 0.035 | 3.0427 | 0.0063 | No |
| | | 12 | 24 | 0.140 | 0.026 | 3.0570 | 0.0050 | Yes |
| | | 13 | 25 | −0.030 | 0.012 | 3.0291 | 0.0033 | Yes |
| | | 14 | 26 | 0 | – | 3.0340 | 0.0026 | Yes |
| | | 15 | 27 | −0.008 | 0.011 | 3.0327 | 0.0031 | No |
| | | 16 | 28 | 0.216 | 0.029 | 3.0694 | 0.0054 | No |
| | | 17 | 29 | 0.256 | 0.037 | 3.0759 | 0.0065 | No |
| | | 18 | 30 | 0.473 | 0.057 | 3.1110 | 0.0095 | No |
| | | 19 | 31 | 0.710 | 0.081 | 3.1488 | 0.0131 | No |
| | | 20 | 32 | 0.948 | 0.102 | 3.1864 | 0.0162 | No |
| 19 | K | 17 | 36 | −0.16 | 0.10 | 3.4115 | 0.0148 | No |
| | | 18 | 37 | −0.08 | 0.07 | 3.4232 | 0.0104 | No |
| | | 19 | 38 | −0.089 | 0.025 | 3.4219 | 0.0041 | Yes |
| | | 20 | 39 | 0 | – | 3.4349 | 0.0019 | Yes |
| | | 21 | 40 | 0.016 | 0.023 | 3.4372 | 0.0038 | Yes |
| | | 22 | 41 | 0.117 | 0.044 | 3.4519 | 0.0066 | Yes |
| | | 23 | 42 | 0.111 | 0.065 | 3.4510 | 0.0096 | Yes |
| | | 24 | 43 | 0.129 | 0.084 | 3.4536 | 0.0123 | Yes |
| | | 25 | 44 | 0.122 | 0.102 | 3.4526 | 0.0149 | Yes |
| | | 26 | 45 | 0.151 | 0.119 | 3.4568 | 0.0173 | Yes |
| | | 27 | 46 | 0.092 | 0.136 | 3.4483 | 0.0198 | Yes |
| | | 28 | 47 | 0.079 | 0.152 | 3.4464 | 0.0221 | Yes |
| | | 29 | 48 | 0.264 | 0.168 | 3.4731 | 0.0243 | No |
| | | 30 | 49 | 0.420 | 0.182 | 3.4955 | 0.0261 | No |
| | | 31 | 50 | 0.513 | 0.197 | 3.5088 | 0.0281 | No |
| | | 32 | 51 | 0.62 | 0.22 | 3.5240 | 0.0313 | No |
| 20 | Ca | 16 | 36 | −0.196 | 0.027 | 3.4493 | 0.0044 | No |
| | | 17 | 37 | −0.205 | 0.023 | 3.4480 | 0.0038 | No |
| | | 18 | 38 | −0.0797 | 0.0064 | 3.4661 | 0.0021 | No |
| | | 19 | 39 | −0.1060 | 0.0064 | 3.4623 | 0.0021 | Yes |
| | | 20 | 40 | 0 | – | 3.4776 | 0.0019 | Yes |
| | | 23 | 43 | 0.114 | 0.009 | 3.4940 | 0.0023 | Yes |
| | | 24 | 44 | 0.288 | 0.007 | 3.5188 | 0.0021 | Yes |
| | | 25 | 45 | 0.125 | 0.009 | 3.4955 | 0.0023 | Yes |
| | | 26 | 46 | 0.125 | 0.009 | 3.4955 | 0.0023 | Yes |
| | | 27 | 47 | 0.002 | 0.010 | 3.4779 | 0.0024 | Yes |
| | | 28 | 48 | 0.001 | 0.011 | 3.4777 | 0.0025 | Yes |
| | | 29 | 49 | 0.098 | 0.013 | 3.4917 | 0.0027 | No |
| | | 30 | 50 | 0.291 | 0.013 | 3.5192 | 0.0026 | Yes |
| | | 31 | 51 | 0.392 | 0.014 | 3.5335 | 0.0027 | No |
| | | 32 | 52 | 0.531 | 0.016 | 3.5531 | 0.0029 | No |
| 25 | Mn | 25 | 50 | 0.168 | 0.054 | 3.7283 | 0.0076 | Yes |
| | | 26 | 51 | 0.065 | 0.067 | 3.7145 | 0.0093 | Yes |
| | | 27 | 52 | −0.226 | 0.034 | 3.6751 | 0.0051 | Yes |
| | | 28 | 53 | −0.285 | 0.022 | 3.6670 | 0.0037 | Yes |
| | | 29 | 54 | −0.166 | 0.013 | 3.6832 | 0.0028 | Yes |
| | | 30 | 55 | 0 | – | 3.7057 | 0.0022 | Yes |
| | | 31 | 56 | 0.053 | 0.016 | 3.7128 | 0.0031 | Yes |
| | | 32 | 57 | 0.202 | 0.022 | 3.7329 | 0.0037 | No |
| | | 33 | 58 | 0.234 | 0.030 | 3.7371 | 0.0046 | No |
| | | 34 | 59 | 0.297 | 0.043 | 3.7456 | 0.0061 | No |
| | | 35 | 60 | 0.329 | 0.048 | 3.7498 | 0.0068 | No |
| | | 36 | 61 | 0.504 | 0.055 | 3.7731 | 0.0076 | No |
| | | 37 | 62 | 0.615 | 0.064 | 3.7878 | 0.0087 | No |
| | | 38 | 63 | 0.704 | 0.070 | 3.7995 | 0.0095 | No |
| | | 39 | 64 | 0.873 | 0.078 | 3.8217 | 0.0104 | No |
| 26 | Fe | 26 | 52 | 0.282 | 0.075 | 3.7313 | 0.0102 | No |
| | | 27 | 53 | 0.097 | 0.038 | 3.7064 | 0.0055 | No |
| | | 28 | 54 | 0 | – | 3.6933 | 0.0019 | Yes |
| 28 | Ni | 30 | 58 | −0.275 | 0.007 | 3.7756 | 0.0019 | Yes |
| | | 32 | 60 | 0 | – | 3.8118 | 0.0016 | Yes |
| | | 33 | 61 | 0.083 | 0.005 | 3.8227 | 0.0017 | Yes |
| | | 34 | 62 | 0.223 | 0.005 | 3.8409 | 0.0017 | Yes |
| | | 36 | 64 | 0.368 | 0.009 | 3.8598 | 0.0020 | Yes |

(continued on next page)

Table 1 (continued)

| Z | el. | N | A | $\delta\langle r^2 \rangle$ (fm 2) | $\Delta\delta\langle r^2 \rangle$ (fm 2) | r_c (fm) | Δr_c (fm) | In Ref. [1] |
|----|-----|-----------|-----------|--|--|---------------|-------------------|-------------|
| 29 | Cu | 40 | 68 | 0.620 | 0.021 | 3.8923 | 0.0031 | No |
| | | 29 | 58 | -0.833 | 0.092 | 3.7940 | 0.0122 | No |
| | | 30 | 59 | -0.635 | 0.072 | 3.8200 | 0.0095 | No |
| | | 31 | 60 | -0.511 | 0.058 | 3.8362 | 0.0077 | No |
| | | 32 | 61 | -0.359 | 0.041 | 3.8559 | 0.0055 | No |
| | | 33 | 62 | -0.293 | 0.034 | 3.8645 | 0.0046 | No |
| | | 34 | 63 | -0.148 | 0.017 | 3.8832 | 0.0026 | Yes |
| | | 35 | 64 | -0.116 | 0.014 | 3.8873 | 0.0023 | No |
| | | 36 | 65 | 0 | - | 3.9022 | 0.0014 | Yes |
| | | 37 | 66 | 0.033 | 0.013 | 3.9064 | 0.0022 | No |
| | | 38 | 67 | 0.115 | 0.019 | 3.9169 | 0.0028 | No |
| | | 39 | 68 | 0.133 | 0.032 | 3.9192 | 0.0043 | No |
| | | 40 | 69 | 0.238 | 0.035 | 3.9326 | 0.0047 | No |
| | | 41 | 70 | 0.271 | 0.045 | 3.9368 | 0.0059 | No |
| | | 42 | 71 | 0.407 | 0.046 | 3.9540 | 0.0060 | No |
| | | 43 | 72 | 0.429 | 0.056 | 3.9568 | 0.0072 | No |
| | | 44 | 73 | 0.523 | 0.060 | 3.9686 | 0.0077 | No |
| | | 45 | 74 | 0.505 | 0.075 | 3.9664 | 0.0096 | No |
| | | 46 | 75 | 0.546 | 0.083 | 3.9715 | 0.0105 | No |
| | | 47 | 76 | 0.58 | 0.11 | 3.9758 | 0.0139 | No |
| | | 48 | 77 | 0.59 | 0.12 | 3.9771 | 0.0151 | No |
| | | 49 | 78 | 0.58 | 0.12 | 3.9758 | 0.0152 | No |
| 30 | Zn | 32 | 62 | -0.493 | 0.053 | 3.9031 | 0.0069 | No |
| | | 33 | 63 | -0.389 | 0.044 | 3.9164 | 0.0058 | No |
| | | 34 | 64 | -0.279 | 0.035 | 3.9305 | 0.0047 | Yes |
| | | 35 | 65 | -0.257 | 0.026 | 3.9333 | 0.0036 | No |
| | | 36 | 66 | -0.121 | 0.017 | 3.9505 | 0.0026 | Yes |
| | | 37 | 67 | -0.089 | 0.010 | 3.9546 | 0.0019 | Yes |
| | | 38 | 68 | 0 | - | 3.9658 | 0.0014 | Yes |
| | | 39 | 69 | 0.026 | 0.011 | 3.9691 | 0.0020 | No |
| | | 40 | 70 | 0.142 | 0.016 | 3.9837 | 0.0024 | Yes |
| | | 41 | 71 | 0.227 | 0.024 | 3.9943 | 0.0033 | No |
| | | 42 | 72 | 0.292 | 0.031 | 4.0024 | 0.0041 | No |
| | | 43 | 73 | 0.318 | 0.038 | 4.0057 | 0.0049 | No |
| | | 44 | 74 | 0.375 | 0.045 | 4.0128 | 0.0058 | No |
| | | 45 | 75 | 0.349 | 0.052 | 4.0096 | 0.0066 | No |
| | | 46 | 76 | 0.421 | 0.058 | 4.0185 | 0.0073 | No |
| | | 47 | 77 | 0.440 | 0.065 | 4.0209 | 0.0082 | No |
| | | 48 | 78 | 0.474 | 0.071 | 4.0251 | 0.0089 | No |
| | | 49 | 79 | 0.461 | 0.078 | 4.0235 | 0.0098 | No |
| | | 50 | 80 | 0.465 | 0.085 | 4.0240 | 0.0107 | No |
| 31 | Ga | 32 | 63 | -0.643 | 0.136 | 3.9308 | 0.0174 | No |
| | | 33 | 64 | -0.579 | 0.121 | 3.9390 | 0.0155 | No |
| | | 34 | 65 | -0.422 | 0.015 | 3.9589 | 0.0026 | No |
| | | 35 | 66 | -0.329 | 0.077 | 3.9706 | 0.0099 | No |
| | | 36 | 67 | -0.252 | 0.031 | 3.9803 | 0.0043 | No |
| | | 37 | 68 | -0.214 | 0.050 | 3.9850 | 0.0065 | No |
| | | 38 | 69 | -0.116 | 0.030 | 3.9973 | 0.0042 | Yes |
| | | 39 | 70 | -0.096 | 0.021 | 3.9998 | 0.0032 | No |
| | | 40 | 71 | 0 | - | 4.0118 | 0.0018 | Yes |
| | | 41 | 72 | 0.161 | 0.028 | 4.0318 | 0.0039 | No |
| | | 42 | 73 | 0.243 | 0.043 | 4.0420 | 0.0056 | No |
| | | 43 | 74 | 0.223 | 0.046 | 4.0395 | 0.0060 | No |
| | | 44 | 75 | 0.285 | 0.059 | 4.0472 | 0.0075 | No |
| | | 45 | 76 | 0.276 | 0.065 | 4.0461 | 0.0082 | No |
| | | 46 | 77 | 0.308 | 0.075 | 4.0500 | 0.0094 | No |
| | | 47 | 78 | 0.270 | 0.079 | 4.0453 | 0.0099 | No |
| | | 48 | 79 | 0.290 | 0.088 | 4.0478 | 0.0110 | No |
| | | 49 | 80 | 0.242 | 0.092 | 4.0418 | 0.0115 | No |
| | | 50 | 81 | 0.242 | 0.099 | 4.0418 | 0.0124 | No |
| | | 51 | 82 | 0.447 | 0.123 | 4.0671 | 0.0152 | No |
| 37 | Rb | 37 | 74 | -0.045 | 0.143 | 4.1935 | 0.0172 | No |
| | | 38 | 75 | 0.266 | 0.129 | 4.2305 | 0.0154 | No |
| | | 39 | 76 | 0.298 | 0.116 | 4.2342 | 0.0139 | Yes |
| | | 50 | 87 | 0 | - | 4.1989 | 0.0021 | Yes |
| 47 | Ag | 50 | 97 | -1.29 | 0.22 | 4.4202 | 0.0250 | No |
| | | 51 | 98 | -1.01 | 0.18 | 4.4518 | 0.0204 | No |
| | | 52 | 99 | -0.91 | 0.14 | 4.4630 | 0.0159 | No |
| | | 53 | 100 | -0.83 | 0.12 | 4.4719 | 0.0137 | No |
| | | 54 | 101 | -0.70 | 0.12 | 4.4865 | 0.0136 | Yes |
| | | 60 | 107 | -0.15 | 0.05 | 4.5473 | 0.0060 | Yes |

(continued on next page)

Table 1 (continued)

| Z | el. | N | A | $\delta\langle r^2 \rangle$ (fm 2) | $\Delta\delta\langle r^2 \rangle$ (fm 2) | r_c (fm) | Δr_c (fm) | In Ref. [1] |
|----|-----|------------|------------|--|--|---------------|-------------------|-------------|
| | | 62 | 109 | 0 | - | 4.5638 | 0.0025 | Yes |
| 48 | Cd | 52 | 100 | -1.421 | 0.044 | 4.4519 | 0.0055 | No |
| | | 53 | 101 | -1.307 | 0.041 | 4.4647 | 0.0052 | No |
| | | 54 | 102 | -1.144 | 0.026 | 4.4829 | 0.0037 | Yes |
| | | 55 | 103 | -1.046 | 0.025 | 4.4938 | 0.0036 | Yes |
| | | 56 | 104 | -0.904 | 0.017 | 4.5096 | 0.0030 | Yes |
| | | 57 | 105 | -0.823 | 0.017 | 4.5185 | 0.0030 | Yes |
| | | 58 | 106 | -0.695 | 0.013 | 4.5327 | 0.0027 | Yes |
| | | 59 | 107 | -0.625 | 0.013 | 4.5404 | 0.0027 | Yes |
| | | 60 | 108 | -0.510 | 0.010 | 4.5530 | 0.0026 | Yes |
| | | 61 | 109 | -0.445 | 0.010 | 4.5602 | 0.0026 | Yes |
| | | 62 | 110 | -0.334 | 0.007 | 4.5723 | 0.0024 | Yes |
| | | 63 | 111 | -0.288 | 0.012 | 4.5773 | 0.0027 | Yes |
| | | 64 | 112 | -0.159 | 0.005 | 4.5914 | 0.0024 | Yes |
| | | 65 | 113 | -0.114 | 0.010 | 4.5963 | 0.0025 | Yes |
| | | 66 | 114 | 0 | - | 4.6087 | 0.0023 | Yes |
| | | 67 | 115 | 0.043 | 0.012 | 4.6134 | 0.0026 | Yes |
| | | 68 | 116 | 0.134 | 0.009 | 4.6232 | 0.0025 | Yes |
| | | 69 | 117 | 0.171 | 0.021 | 4.6272 | 0.0032 | Yes |
| | | 70 | 118 | 0.243 | 0.023 | 4.6350 | 0.0034 | Yes |
| | | 71 | 119 | 0.283 | 0.034 | 4.6393 | 0.0043 | No |
| | | 72 | 120 | 0.342 | 0.038 | 4.6457 | 0.0047 | Yes |
| | | 73 | 121 | 0.375 | 0.050 | 4.6492 | 0.0058 | No |
| | | 74 | 122 | 0.431 | 0.055 | 4.6552 | 0.0063 | No |
| | | 75 | 123 | 0.457 | 0.068 | 4.6580 | 0.0076 | No |
| | | 76 | 124 | 0.510 | 0.073 | 4.6637 | 0.0081 | No |
| | | 77 | 125 | 0.533 | 0.087 | 4.6662 | 0.0096 | No |
| | | 78 | 126 | 0.585 | 0.091 | 4.6717 | 0.0100 | No |
| | | 79 | 127 | 0.599 | 0.107 | 4.6732 | 0.0117 | No |
| | | 80 | 128 | 0.660 | 0.108 | 4.6798 | 0.0118 | No |
| | | 81 | 129 | 0.638 | 0.134 | 4.6774 | 0.0145 | No |
| | | 82 | 130 | 0.705 | 0.133 | 4.6846 | 0.0144 | No |
| 50 | Sn | 58 | 108 | -1.081 | 0.018 | 4.5579 | 0.0022 | Yes |
| | | 60 | 110 | -0.907 | 0.011 | 4.5770 | 0.0016 | Yes |
| | | 62 | 112 | -0.747 | 0.007 | 4.5944 | 0.0013 | Yes |
| | | 64 | 114 | -0.605 | 0.005 | 4.6098 | 0.0012 | Yes |
| | | 66 | 116 | -0.461 | 0.004 | 4.6254 | 0.0011 | Yes |
| | | 68 | 118 | -0.324 | 0.005 | 4.6402 | 0.0011 | Yes |
| | | 70 | 120 | -0.206 | 0.005 | 4.6529 | 0.0011 | Yes |
| | | 72 | 122 | -0.097 | 0.003 | 4.6646 | 0.0011 | Yes |
| | | 74 | 124 | 0 | - | 4.6750 | 0.0010 | Yes |
| | | 76 | 126 | 0.089 | 0.004 | 4.6845 | 0.0011 | Yes |
| | | 78 | 128 | 0.165 | 0.010 | 4.6926 | 0.0015 | Yes |
| | | 80 | 130 | 0.240 | 0.015 | 4.7006 | 0.0019 | Yes |
| | | 82 | 132 | 0.307 | 0.021 | 4.7077 | 0.0024 | Yes |
| | | 84 | 134 | 0.533 | 0.010 | 4.7317 | 0.0014 | No |
| 70 | Yb | 105 | 175 | -0.0543 | 0.0002 | 5.3164 | 0.0062 | Yes |
| | | 106 | 176 | 0 | - | 5.3215 | 0.0062 | Yes |
| | | 107 | 177 | 0.0259 | 0.0001 | 5.3239 | 0.0062 | No |
| 80 | Hg | 97 | 177 | -1.067 | 0.079 | 5.3474 | 0.0080 | No |
| | | 98 | 178 | -0.968 | 0.072 | 5.3567 | 0.0074 | No |
| | | 99 | 179 | -0.905 | 0.071 | 5.3626 | 0.0073 | No |
| | | 100 | 180 | -0.808 | 0.061 | 5.3716 | 0.0065 | No |
| | | 101 | 181 | -0.111 | 0.013 | 5.4361 | 0.0033 | Yes |
| | | 102 | 182 | -0.653 | 0.049 | 5.3860 | 0.0055 | Yes |
| | | 103 | 183 | -0.065 | 0.009 | 5.4403 | 0.0032 | Yes |
| | | 104 | 184 | -0.542 | 0.041 | 5.3963 | 0.0049 | Yes |
| | | 105 | 185 | -0.069 | 0.010 | 5.4400 | 0.0032 | Yes |
| | | 118 | 198 | 0 | - | 5.4463 | 0.0031 | Yes |
| | | 98 | 179 | -1.274 | 0.094 | 5.3583 | 0.0092 | No |
| 81 | Tl | 99 | 180 | -1.254 | 0.091 | 5.3602 | 0.0089 | No |
| | | 100 | 181 | -1.174 | 0.084 | 5.3676 | 0.0083 | No |
| | | 102 | 183 | -1.033 | 0.074 | 5.3786 | 0.0074 | No |
| | | 104 | 185 | -0.938 | 0.078 | 5.3896 | 0.0077 | No |
| | | 109 | 190 | -0.7063 | 0.0490 | 5.4110 | 0.0052 | Yes |
| | | 110 | 191 | -0.6544 | 0.0460 | 5.4158 | 0.0050 | Yes |
| | | 111 | 192 | -0.6296 | 0.0440 | 5.4181 | 0.0048 | Yes |
| | | 112 | 193 | -0.5716 | 0.0400 | 5.4235 | 0.0045 | Yes |
| | | 113 | 194 | -0.5551 | 0.0063 | 5.4250 | 0.0027 | Yes |
| | | 114 | 195 | -0.4820 | 0.0347 | 5.4317 | 0.0041 | Yes |
| | | 115 | 196 | -0.4795 | 0.0340 | 5.4319 | 0.0041 | Yes |

(continued on next page)

Table 1 (continued)

| Z | el. | N | A | $\delta\langle r^2 \rangle$ (fm 2) | $\Delta\delta\langle r^2 \rangle$ (fm 2) | r_c (fm) | Δr_c (fm) | In Ref. [1] |
|----|-----|------------|------------|--|--|---------------|-------------------|-------------|
| 82 | Te | 116 | 197 | -0.4119 | 0.0299 | 5.4382 | 0.0038 | Yes |
| | | 117 | 198 | -0.4035 | 0.0289 | 5.4389 | 0.0037 | Yes |
| | | 118 | 199 | -0.3116 | 0.0231 | 5.4474 | 0.0034 | Yes |
| | | 119 | 200 | -0.2979 | 0.0222 | 5.4486 | 0.0033 | Yes |
| | | 120 | 201 | -0.2077 | 0.0150 | 5.4569 | 0.0030 | Yes |
| | | 121 | 202 | -0.1834 | 0.0148 | 5.4591 | 0.0029 | Yes |
| | | 122 | 203 | -0.1032 | 0.0070 | 5.4665 | 0.0027 | Yes |
| | | 123 | 204 | -0.0635 | 0.0081 | 5.4701 | 0.0027 | Yes |
| | | 124 | 205 | 0 | - | 5.4759 | 0.0026 | Yes |
| | | 126 | 207 | 0.1048 | 0.0070 | 5.4855 | 0.0027 | Yes |
| | | 127 | 208 | 0.183 | 0.019 | 5.4926 | 0.0031 | Yes |
| | | 128 | 209 | 0 | - | 5.5211 | 0.0026 | Yes |
| 83 | Bi | 128 | 211 | 0.221 | 0.017 | 5.5411 | 0.0031 | No |
| | | 130 | 213 | 0.422 | 0.029 | 5.5592 | 0.0038 | Yes |
| 84 | Po | 109 | 193 | -0.576 | 0.013 | 5.5185 | 0.0178 | No |
| | | 111 | 195 | -0.604 | 0.013 | 5.5159 | 0.0178 | No |
| | | 113 | 197 | -0.657 | 0.013 | 5.5111 | 0.0178 | No |
| | | 115 | 199 | -0.644 | 0.013 | 5.5123 | 0.0178 | No |
| | | 117 | 201 | -0.510 | 0.013 | 5.5244 | 0.0178 | No |
| | | 119 | 203 | -0.425 | 0.013 | 5.5321 | 0.0178 | No |
| | | 125 | 209 | -0.0813 | 0.0100 | 5.5631 | 0.0176 | Yes |
| | | 125 | 210 | 0 | - | 5.5704 | 0.0176 | Yes |
| | | 127 | 211 | 0.104 | 0.010 | 5.5797 | 0.0176 | No |
| | | 133 | 217 | 0.821 | 0.018 | 5.6436 | 0.0174 | No |
| | | 134 | 218 | 0.948 | 0.013 | 5.6549 | 0.0174 | Yes |
| 87 | Fr | 115 | 202 | -1.596 | 0.018 | 5.5367 | 0.0182 | No |
| | | 116 | 203 | -1.530 | 0.018 | 5.5427 | 0.0182 | No |
| | | 117 | 204 | -1.571 | 0.018 | 5.5390 | 0.0183 | No |
| | | 118 | 205 | -1.475 | 0.017 | 5.5476 | 0.0182 | No |
| | | 119 | 206 | -1.465 | 0.017 | 5.5485 | 0.0182 | No |
| | | 134 | 221 | 0 | - | 5.6790 | 0.0177 | Yes |
| 88 | Ra | 126 | 214 | 0 | - | 5.6079 | 0.0177 | Yes |
| | | 134 | 222 | 1.0449 | 0.0524 | 5.7003 | 0.0180 | Yes |
| | | 125 | 223 | 1.1708 | 0.0587 | 5.7113 | 0.0181 | Yes |
| | | 136 | 224 | 1.2680 | 0.0636 | 5.7198 | 0.0182 | Yes |
| | | 137 | 225 | 1.4041 | 0.0704 | 5.7317 | 0.0184 | Yes |
| | | 138 | 226 | 1.4858 | 0.0745 | 5.7388 | 0.0185 | Yes |
| | | 139 | 227 | 1.5871 | 0.0796 | 5.7477 | 0.0186 | Yes |
| | | 140 | 228 | 1.6980 | 0.0852 | 5.7573 | 0.0188 | Yes |
| | | 141 | 229 | 1.8102 | 0.0908 | 5.7670 | 0.0189 | Yes |
| | | 142 | 230 | 1.9435 | 0.0975 | 5.7786 | 0.0191 | Yes |
| | | 143 | 231 | 2.0177 | 0.1012 | 5.7850 | 0.0193 | No |
| | | 144 | 232 | 2.1589 | 0.1083 | 5.7972 | 0.0195 | Yes |
| | | 145 | 233 | 2.225 | 0.115 | 5.8029 | 0.0198 | No |